



TR&T Update

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TR&T Funding Profile



- \$6.0M in 2000
- \$5.2M in 2001
- \$5.3M in 2002
- \$4.3M in 2003
- \$6.3M in 2004
- \$6.5M in 2005 +
 - \$2.0M for Strategic Capability In partnership with NSF & AFSOR

Larger awards to consortia

- Earth–Moon–Mars Radiation Model
- Comprehensive Magnetospherelonosphere Model
- A Time-Dependent 3D Model of the Corona and Ambient Solar Wind

Total: 150 grants, inter-agency transfers, contracts, RTOPs





Solar

 The magnetic field topology connecting the photosphere to the corona.

Heliosphere

 The propagation of the background solar wind flow and superimposed disturbances through the heliosphere

Magnetosphere

 The generation and decay of the Earth's radiation belts as a function of geomagnetic and solar wind conditions.

Ionosphere

 The geophysical conditions favoring the development of low- and mid-latitude scintillations in the Earth's ionosphere.

ITM

• The effects of varying solar EUV radiation on the Earth's ionosphere and atmosphere.

Solar-Climate

 The relationship between solar irradiance and cosmogenic proxies for long term solar activity.



Team Approach



- FY04-05--Implemented teams for each focus topic
- LWS Science Center MoU with NCAR

AGU Special Sessions

- Dec 2006
- May 2007

Workshop Sep 2007

Team Meetings

- Liemohm GEM Jun 2006
- Richmond NCAR Sep 2006
- Fuller-Rowell NCAR Sep 2006
- Zurbuchen NCAR Nov 2006
- Liemohm AGU Dec 2006





Solar *Mikic*

 Determine the solar origins of the plasma and magnetic flux observed in an Interplanetary Coronal Mass Ejection.

Solar- Heliosphere Zurbuchen

• Determine the topology and evolution of the open magnetic field of the Sun connecting the photosphere through the corona to the heliosphere.

Solar- Heliosphere Desai

• Relate solar-energetic particles to their origin at the sun and inner heliosphere.

Magnetosphere Liemohn

 Determine the mechanisms responsible for the formation and loss of new radiation belts in the slot region in response to geo-effective solar wind structures.

Thermosphere *Richmond*

 Quantify the response of thermospheric density and composition to solar and high latitude forcing.

Solar-Climate

 Quantify the sensitivity of regional and global climate to solar forcing in the full context of the interactive climate system.





Heliosphere Lee

 Shock acceleration of SEPs by interplanetay CMEs

Heliosphere *Miralles*

Mechanism for solar wind heating and acceleration

Magnetosphere *Thomsen*

Solar wind plasma entry and transport in the magnetosphere

Ionosphere Fuller-Rowell

 Storm effects on global electrodynamics and middle and low latitude ionosphere

Solar-Climate

 Atmospheric abundance of greenhouse gases and dynamics of upper atmosphere





 Predict Emergence of Solar Active Regions Before they are Visible

Solar-Heliosphere

 Understand how Flares Accelerate Particles near the Sun (i.e., through shocks and/or Reconnection) and how they Contribute to Large SEP Events

Magnetosphere - lonosphere

 Effects of Ionospheric-Magnetospheric Plasma Redistribution on Storms

ITM

 Investigate the Global Distribution, Sources and Effects of Large Electron Density Gradients at Middle and Low Latitudes

Solar-Climate

Solar Origins of Irradiance Variations



New Collaborations



Collaborations > Sum of Individual Efforts

